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Condensation—mass flux connection in warm convective clouds: theory and implications for cloud supersaturation

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Abstract. The study focused on the relationship between Condensation Rate (CR) and the upward/Plus Mass Flux (MFP) in a system of trade wind cumulus clouds simulated by an LES model. The model was initialized with data observed during the RICO field project, and simulated in a $50.0 \times 50.0 \,\mathrm{km}$ horizontal domain.

In our previous study (Kogan, 2021) we showed that a nearly perfect correlation exists between CR and MFP (correlation coefficient R=0.99). As a result, condensation rate can be highly accurately expressed as a linear function of upward mass flux. This LES derived finding was explained using condensation theory and concept of quasi-steady supersaturation. The obtained from the LES model slope of the CR–MFP linear fit was in excellent agreement with its theoretical value (error less than 5%). The theory also showed that the equality between the LES and theoretical values of the slope follows from the equality between supersaturation and its quasi-steady value.

The study results suggest that condensation rates, for a variety of cloud conditions, can be precisely estimated using the single variable–upward mass flux. Possible implications of the results for evaluating supersaturation and degree of non-adiabaticity in clouds are discussed.

1 Introduction

In tropical shallow cumulus clouds latent heat released during water phase transition is an important source of energy driving formation and evolution of cumulus convection. The formulation of phase transition processes (condensation and evaporation) in computer models depends on grid resolution (LES/CRM/NWP) and the method chosen to describe the microphysics (explicit/bin or parameterized/bulk). Obviously, condensation is influenced by microphysics, including aerosols (Kogan and Martin, 1994). The latter effect is often referred to as "convection invigoration". It is, therefore, important to formulate parameterizations for phase transition processes, especially applicable to cloud resolving (CRM) and NWP models.

In our previous study (Kogan, 2021) we found that a strong correlation exists between integral cloud condensation rate (CR) and integral upward mass flux (MFP). At first sight this result should not be surprising, as mass flux has long

been recognized as a major factor affecting cumulus convection (see e.g., Arakawa and Schubert, 1974; Tiedtke, 1989; Suselj et al., 2019). It was also well known that vertical velocity has a strong effect on supersaturation, and, therefore, cloud microphysics (see, e.g., Squires, 1952; Politovich and Cooper, 1988). Nevertheless, the exceptionally high correlation (correlation coefficient R=0.99), and the simple linear relationship between these variables (CR = α_{les} MFP), was quite remarkable. While the analyzed dataset consisted of more than 2000 clouds in a wide range of sizes and at different stages of their evolution, the slope of the linear fit, α_{les} , remained nearly constant. This result was also obtained in a recent study by Grant et al. (2022) who found a similar linear relationship between condensation and vertical velocity in deep convective clouds.

This study provides theoretical formulation of the CR–MFP relationship based on the cloud drop condensational growth equation and the concept of quasi-steady supersaturation.

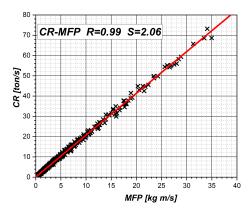


Figure 1. Scatter plot of condensation rates (CR) as a function of upward mass flux (MFP) for clouds in all groups. R is the correlation coefficient; S is the slope of the linear fit.

2 Model and dataset

Our LES model (SAMBM) employs the dynamical core of the System for Atmospheric Modeling (SAM; Khairoutdinov and Randall, 2003) and the Bulk Microphysics (BM; Kogan, 2013) fine-tuned for shallow Cu convection. The observations from the RICO field campaign (vanZanten et al., 2011) were used for initialization of the LES simulations conducted in a rather large $50.0 \times 50.0 \times 4 \, \mathrm{km}^3$ domain $(500 \times 500 \times 100 \, \mathrm{grid} \, \mathrm{points})$.

Over the course of the simulation from 8 to 32 h, we selected 2031 clouds by applying a threshold of liquid water path $> 20 \,\mathrm{g}\,\mathrm{m}^{-2}$. In order to facilitate the analysis of the complex system of clouds at various stages of their development, the dataset was sorted out by cloud top height and divided into four groups G1–G4, each of which condenses approximately equal amount of water vapor per second. The groups G1–G2 represent clouds mostly in the growing stage, while groups G3–G4 contain mature and decaying clouds.

3 LES model results

The analysis of the LES dataset described in Kogan (2021) revealed a remarkably strong correlation between the condensation rate and the upward mass flux. The scatter plot in Fig. 1 shows the CR–MFP dependence for the whole dataset, e.g., for clouds of all sizes and all stages of their evolution.

The data clearly shows a perfect linear relationship (R = 0.99) between the condensation rate ($CR \equiv \frac{dq_l}{dt}$) and upward mass flux ($\rho_a W$) which can be expressed as:

$$\frac{\mathrm{d}q_l}{\mathrm{d}t} = \alpha_{\mathrm{les}}\rho_{\mathrm{a}}W\tag{1}$$

where $\alpha_{\rm les} = 2.06 \times 10^{-6} \ [{\rm m}^{-1}]$. Here liquid water content q_l and air density $\rho_{\rm a}$ are in kg m⁻³, vertical velocity W in m s⁻¹. The subscript "les" signifies that coefficient $\alpha_{\rm les}$ is

obtained from LES data, as opposite to the similar coefficient α_{qs} derived from theory (see below).

The coefficients α_{les} for individual cloud groups slightly differ from the coefficient for the whole dataset (Fig. 2). They are 4 %–5 % higher for small clouds (2.17 for G1 and 2.13 for G2), while lower for larger clouds (2.052 for G3 and 2.049 for G4 clouds). Obviously, using the specific latent heat constant, one can relate formulation (Eq. 1) to the latent heat released during phase transitions.

4 Theoretical formulation of the CR–MFP relationship

The SAMBM model used in our study is a so-called two and half moment model, i.e., it employs three prognostic moments for cloud water variables (cloud water content, q_c , cloud drop concentration, n_c , and cloud integral radius, R_i), as well as two moments for rain water variables (rain water content, q_r , and rain drop concentration, n_r) (Khairoutdinov and Kogan, 1999). The use of an additional cloud moment, the integral radius $R_i = \int rn(r) dr (n(r))$ is the drop size distribution function), is essential, as it allows explicit calculation of condensation rates. Specifically, using the condensational growth equation for cloud drop with radius r

$$\frac{\mathrm{d}r}{\mathrm{d}t} = \frac{G(T, P)S}{r} \tag{2}$$

and having R_i as a prognostic variable, we can explicitly calculate the rate of change of cloud liquid water content:

$$\frac{\mathrm{d}q_l}{\mathrm{d}t} = 4\pi \,\rho_{\mathrm{w}} G(TP) S R_i. \tag{3}$$

Here S is supersaturation, ρ_w and ρ_a are water and air density, respectively. The coefficient G(TP) is a weak function of temperature T and pressure P (e.g., Politovich and Cooper, 1988).

In order to solve Eq. (3), we first calculate changes in temperature T and water vapor content $q_{\rm v}$ due to advection and turbulent mixing. The corresponding intermediate values of T^* and $q_{\rm v}^*$ are used to calculate supersaturation S, and, based on it, condensation/evaporation rates according to Eq. (3). This splitting of dynamical and microphysical terms is analogous to a Lagrangian air parcel model where the two terms in the supersaturation equation (Eq. 4) account for dynamical ascent/descent, as well as latent heat release (e.g., Squires, 1952; Paluch and Knight, 1984).

$$\frac{\mathrm{d}S}{\mathrm{d}t} = A_1 W - A_2 S R_{\mathrm{i}} \tag{4}$$

In Eq. (4) coefficients A_1 and A_2 are weak functions of temperature T and pressure P (see, e.g., Politovich and Cooper, 1988; Pinsky et al., 2013). As was shown in many studies of supersaturation in clouds (e.g., Paluch and Knight, 1984; Cooper, 1989; Korolev and Mazin, 2003; Pinsky et al., 2013;

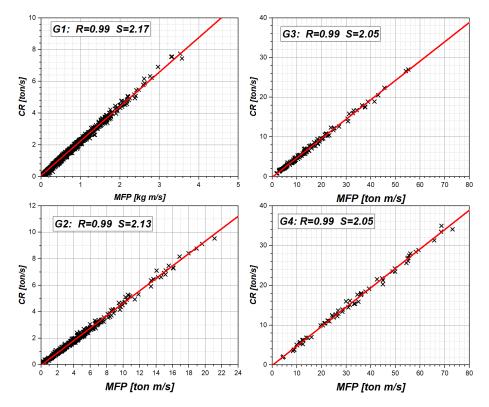


Figure 2. Scatter plots of condensation rates (CR) as a function of upward mass flux (MFP) for clouds in each of the four groups. R is the correlation coefficient; S is the slope of the linear fit.

Siebert and Shaw, 2017), for times larger than phase relaxation time, the Eq. (4) has an asymptotic quasi-steady solution for S:

$$S_{qs} = \frac{A_1 W}{A_2 R_i}. (5)$$

The quasi-steady supersaturation is an important characteristic of the condensation process, as it corresponds to an equilibrium between contributions to the supersaturation equation from dynamical and microphysical processes. It is appropriate, therefore, to measure the actual cloud supersaturation in units of S_{qs} :

$$S = \kappa S_{as}. \tag{6}$$

Under adiabatic conditions for times larger than phase relaxation time (e.g. Paluch and Knight, 1984; Khain et al., 2000; Pinsky et al., 2022) the supersaturation reaches its 'quasisteady' value. In this case the kappa factor $\kappa=1$. In general, due to mixing and entrainment, the value of κ may deviate from 1. Using LES model, the spatial distribution of κ can be evaluated according to Eq. (9) below, and, thus, κ may serve as a measure of entrainment/mixing in various cloud regions.

Substituting Eq. (6) into Eq. (3) and using Eq. (5), we can rewrite Eq. (3) as:

$$\frac{\mathrm{d}q_{l}}{\mathrm{d}t} = \frac{4\pi \rho_{\mathrm{w}} G A_{1}}{A_{2}} \kappa W = \kappa \frac{A_{1}}{A_{3}} \rho_{\mathrm{a}} W = \kappa \alpha_{qs} \rho_{\mathrm{a}} W \tag{7}$$

where coefficient A_3 is expressed through A_2 and G. In Eq. (7) we introduced coefficient

$$\alpha_{qs} = \frac{A_1}{A_3} \tag{8}$$

which is a weak function of temperature T and pressure P (Pinsky et al., 2013). Similar to Eq. (3) formulation (Eq. 7) demonstrates that CR is a linear function of upward mass flux. By comparing the LES derived formulation (Eq. 3) with the theoretical formulation (Eq. 7), one can obtain κ as:

$$\kappa = \alpha_{\rm les}/\alpha_{\rm qs}.\tag{9}$$

In order to evaluate α_{qs} which is a function of temperature T and pressure P, we first calculate at each vertical level the values of T and P as horizontal averages over the cloudy areas, and then obtain α_{qs} using Eq. (8), and expressions for A_1 and A_3 . The values of α_{les} are calculated as slopes of linear fits either to the full LES dataset ($\alpha_{qs} = 2.06$), or to each individual group subset (2.17, 2.13, 2.052, and 2.049 for G1–G4 groups, respectively, see Fig. 2).

The vertical profile of κ is shown in Fig. 3 for two times in cloud system evolution. The left panel refers to an early cloud system formation ($t=8\,\mathrm{h}$) when only small clouds in G1/G2 groups were formed. The right panel refers to $t=24\,\mathrm{h}$ when the cloud system has already been fully developed with clouds present in all groups.

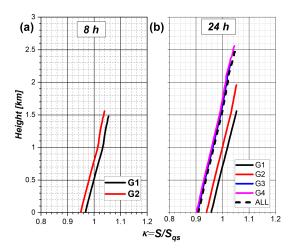


Figure 3. Vertical profile of the ratio of supersaturation to its quasisteady value. (a) The early stage of cloud formation (time = 8 h). (b) the mature stage of cloud system development (time = 24 h). The vertical coordinate is the height above the cloud base.

As Fig. 3 shows, in the early formed clouds at $t = 8 \,\mathrm{h}$ the supersaturation is, on average, smaller by 3 %–5 % than its quasi-steady value (Fig. 3a). This is the result of small values of integral radius R_i at the early stage of cloud formation, which combined with growing updrafts leads to large quasi-steady supersaturations (Eq. 5). At the mature stage R_i is larger, resulting in the decrease of S_{qs} . The decrease in S_{qs} is especially pronounced below cloud tops where updrafts approaching the inversion layer are weak.

While all these factors may qualitatively explain the increase of κ with height, the more accurate profile of κ may be obtained by accounting for vertical variation of α_{les} . The profiles shown in Fig. 3 consider only constant with height integral values of α_{les} ; i.e., averaged over the whole cloud. Our preliminary results from the study of vertically dependent CR and MFP variables, showed that the coefficient α_{les} slightly increases with height. As the result, the qualitative behavior of κ with height will remain the same.

5 Conclusions

- Data from LES simulations of shallow cumulus clouds demonstrated a nearly perfect correlation between Condensation Rate (CR) and upward/Plus Mass Flux (MFP).
- The strong correlation and the linear relationship between these variables, is explained using the condensation theory. The theory also shows that the linear CR–MFP dependence is due to the fact that cloud supersaturations, on average, are equal to their quasi-steady values.
- 3. In cloud regions where, because of entrainment and mixing, supersaturation differs from its quasi-steady

- value, the degree of non-adiabaticity may be evaluated by the κ -factor. The latter is defined as the ratio of supersaturation to its quasi-steady value ($\kappa = S/S_{qs}$), and can be calculated using the slope of the CR–MFP linear fit and relationship (Eq. 9).
- 4. The strong dependence of condensation rates on vertical mass flux, and the simple linear functional relationship between these variables, may be useful in formulation of condensation process in simple conceptual models of cloud topped convective boundary layer.
- 5. Finally, we note that the theoretical formulation of the CR–MFP linear relationship was based on the cloud drop condensational growth equation, and, thus, applicable for local variables. As a result, it may be integrated over limited regions of the clouds, e.g., over horizontal cloud slices to obtain relationships for horizontally averaged variables. Expanding the CR–MFP relationship for horizontally averaged variables which vary in the vertical, may serve as a framework for sub-grid scale latent heat release parameterization.

Code availability. The software code for data analysis was developed by the author and is available upon request.

Data availability. Analysis data is available upon request from the author.

Competing interests. The author has declared that there are no competing interests.

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